



**Pacific  
Northwest**  
NATIONAL LABORATORY

# Cosmogenic Activation Backgrounds

**Richard Saldanha**

2<sup>nd</sup> October 2024  
LRT 2024

U.S. DEPARTMENT OF  
**ENERGY** **BATTELLE**

PNNL is operated by Battelle for the U.S. Department of Energy



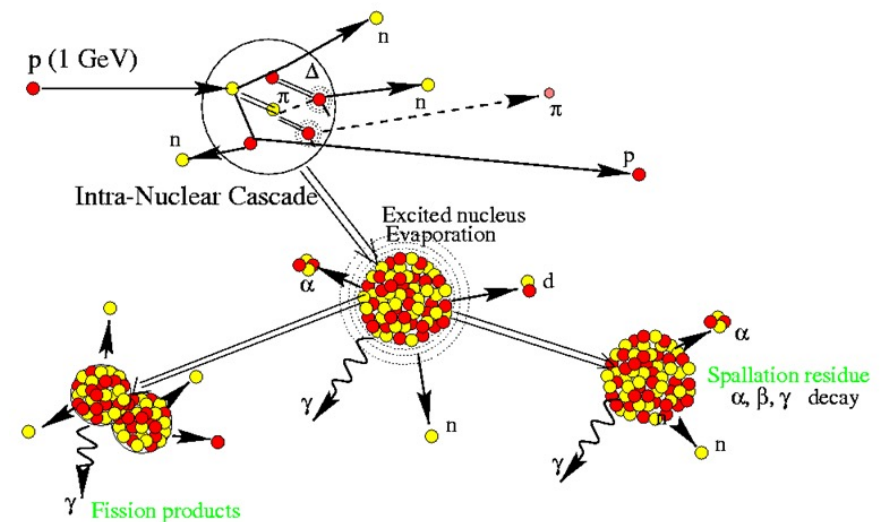
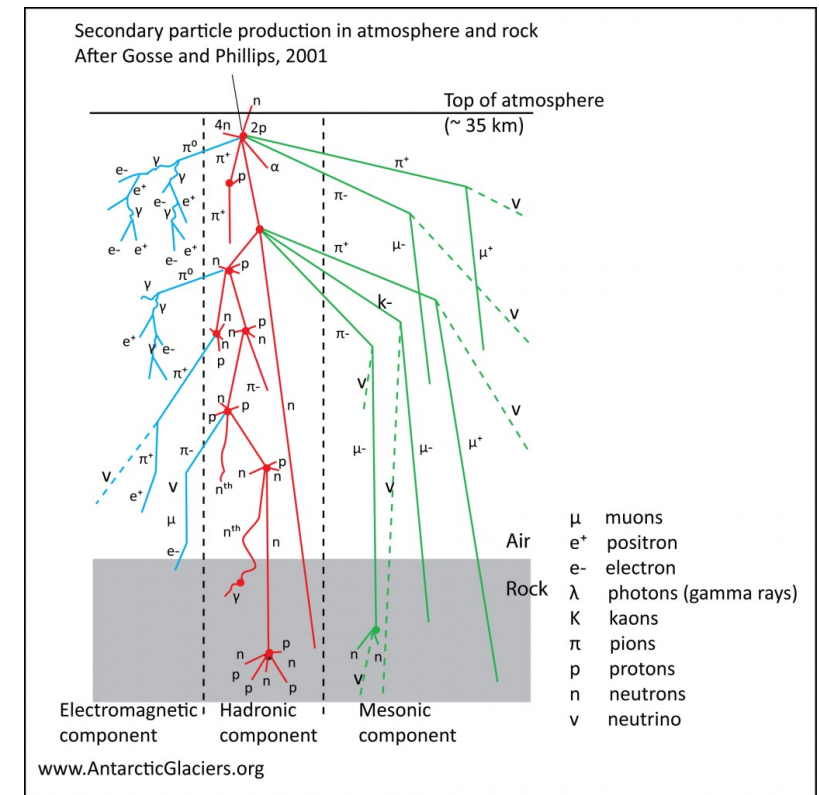
NSF/J. Yang



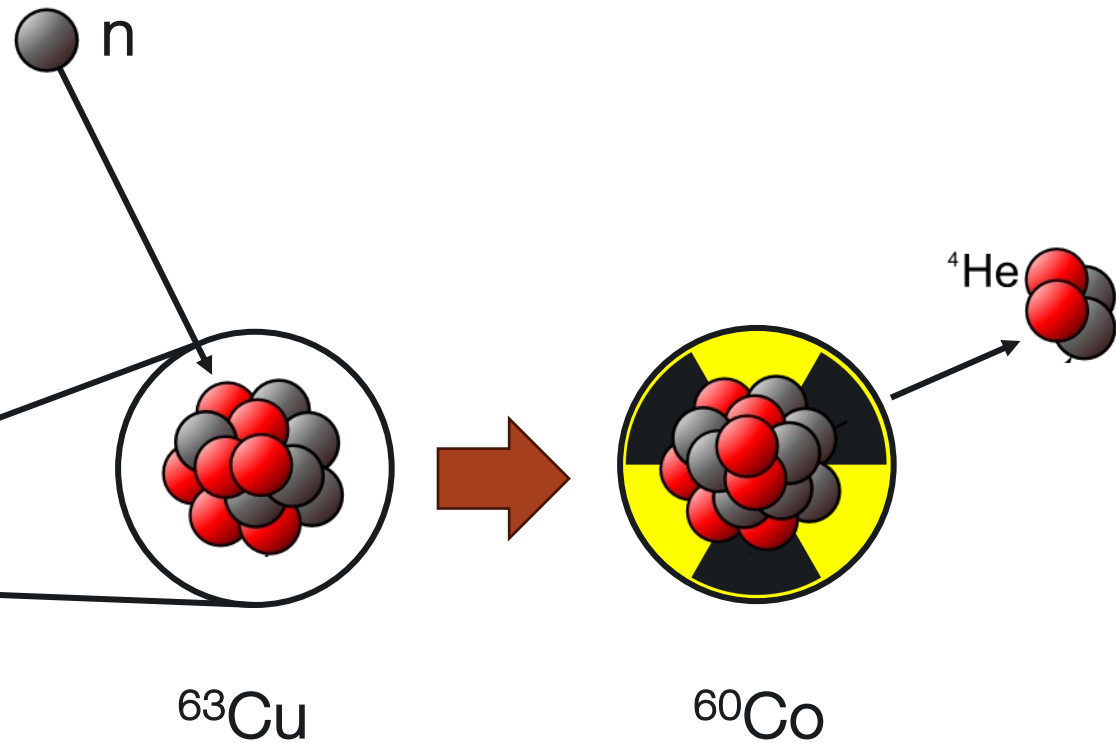
# Cosmogenic Activation

As low-radioactivity experiments get bigger and more sensitive, understanding the production rate of cosmogenic isotopes becomes more and more important

Needed to evaluate the total surface residency time, transportation options, and storage requirements for all low background detector components



# Cosmogenic Activation



Sea-level Production Rate:  
30-90 atoms/kg/day

Sea-level Saturation Rate:  
0.3-1.0 mBq/kg

## Cosmogenic Production Rate

$$P(\vec{r}, t) = \int \Phi(E, \vec{r}, t) \sigma(E) n dE$$

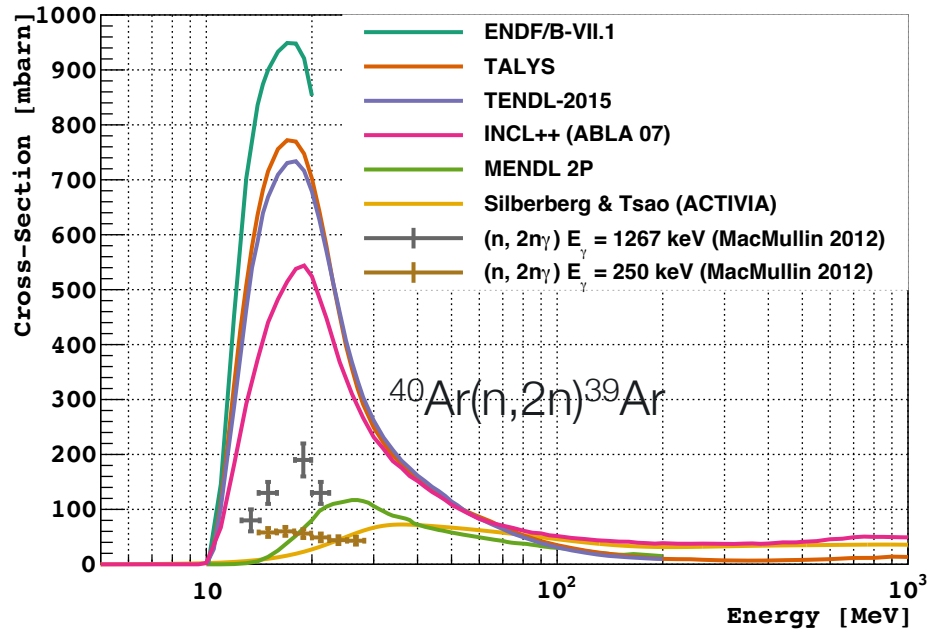
The production rate depends on 3 things

- Cosmogenic particle flux
- Production cross-section
- Target material



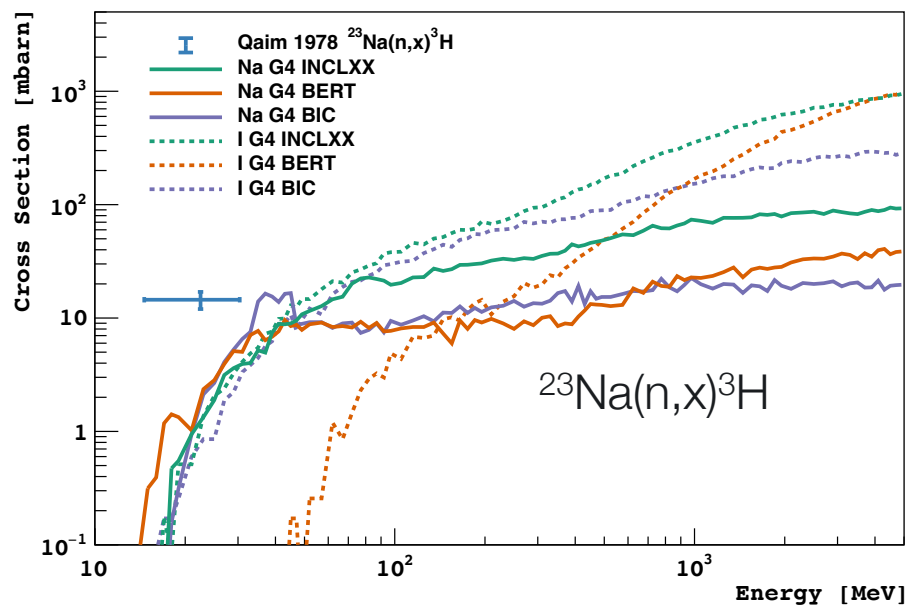
# Cosmogenic Activation Cross-Sections

There are **very few experimental measurements** of key production cross-sections



- Difficult to produce monoenergetic high energy neutron beams
- Products of interest to dark matter and other rare event searches can be low energy and not easy to detect

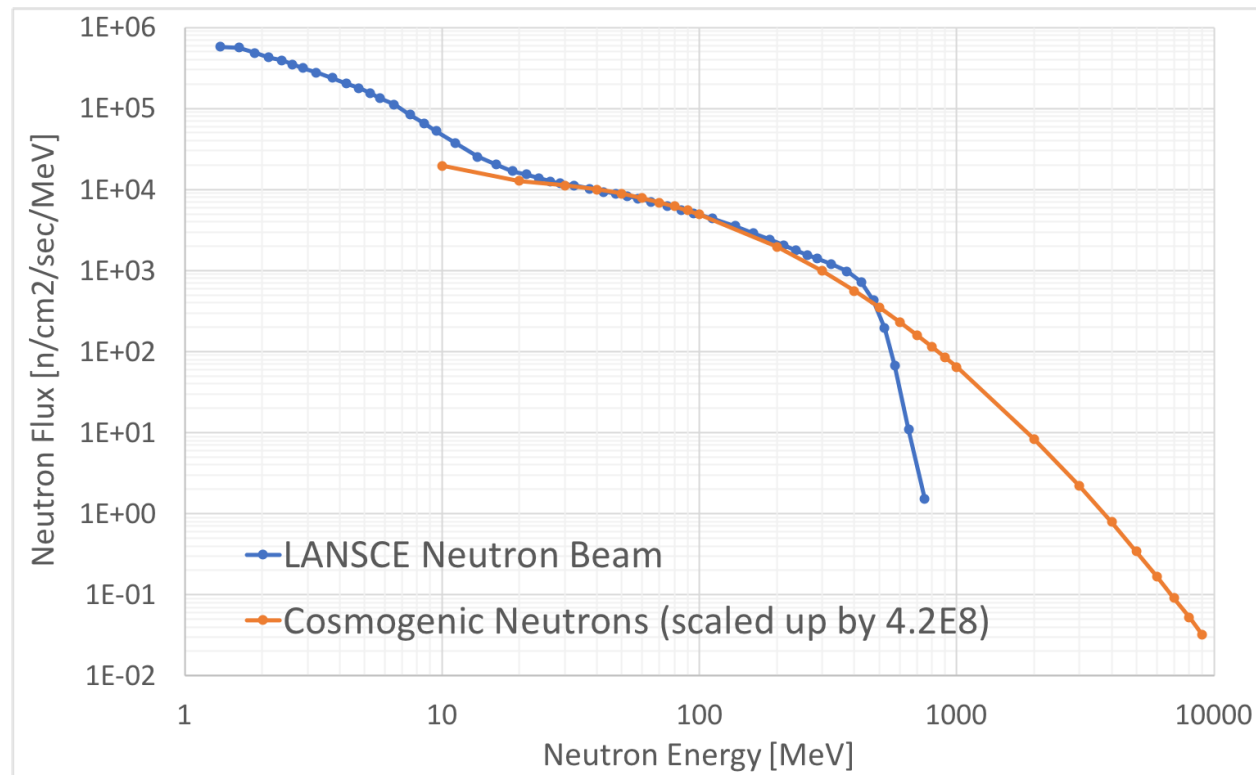
Cross-sections from nuclear data libraries and simulation codes often vary by an order of magnitude or larger



Can we directly measure the production rate through detection of decays?

# LANSCE ICE-HOUSE Neutron Beam

Los Alamos Neutron Science Center (LANSCE) has a neutron beam (4FP30R ICE-HOUSE II) that is very similar in spectral shape to the cosmic ray spectrum



The good agreement in spectral shape from 10–500 MeV allows for low-uncertainty extrapolations to cosmic ray activation rates

The neutron flux is roughly  **$4.2 \times 10^8$  times larger** than the sea-level cosmic neutron flux

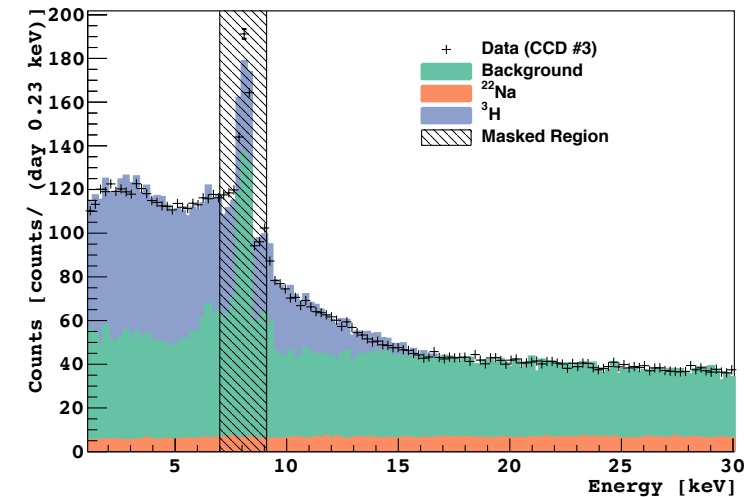
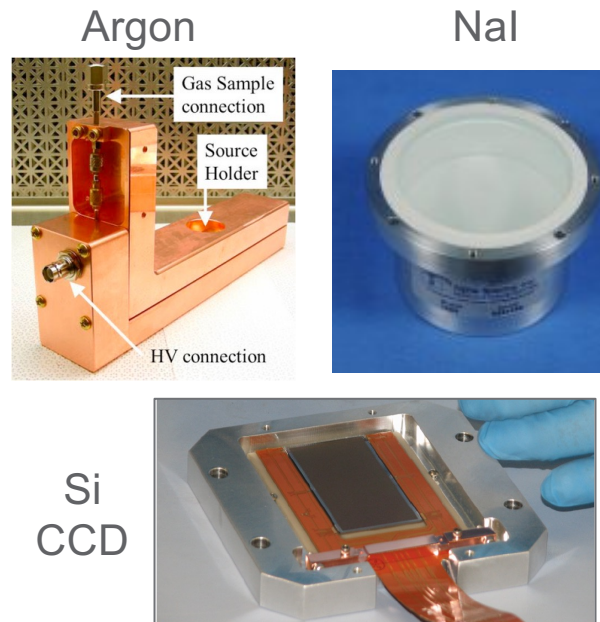
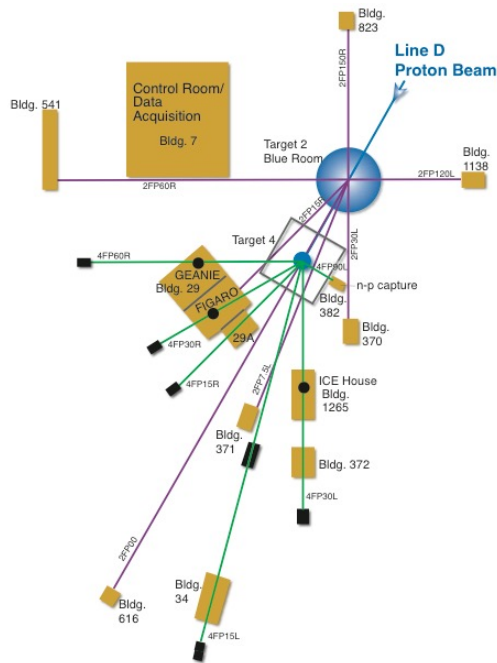
1 second on beam  
~ 13 years on the surface

# Measurement Technique

Use high intensity neutron beam to increase production rate compared to sea-level cosmic rays

Irradiate material of interest and use radiation detection techniques to directly measure decay of activation products

Use **known** neutron fluence, and **measured activity** to extrapolate to natural cosmogenic production rate



# Activation Measurements for Dark Matter Experiments

$^{39}\text{Ar}$ ,  $^{37}\text{Ar}$  in Argon



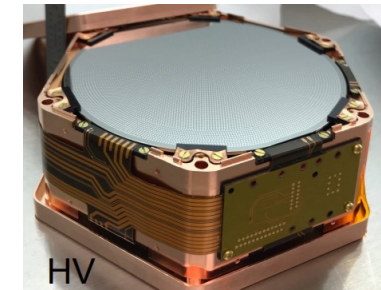
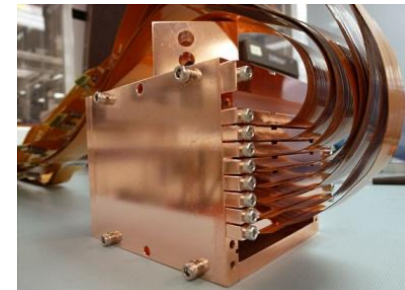
Phys. Rev. C 100, 024608 (2019)  
arXiv:1902.09072



$^3\text{H}$  (Tritium),  $^7\text{Be}$ ,  $^{22}\text{Na}$  in Silicon



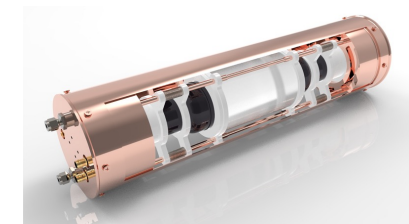
Phys. Rev. D 102, 102006 (2020)  
arXiv:2007.10584



Cosmogenic activation of Sodium Iodide

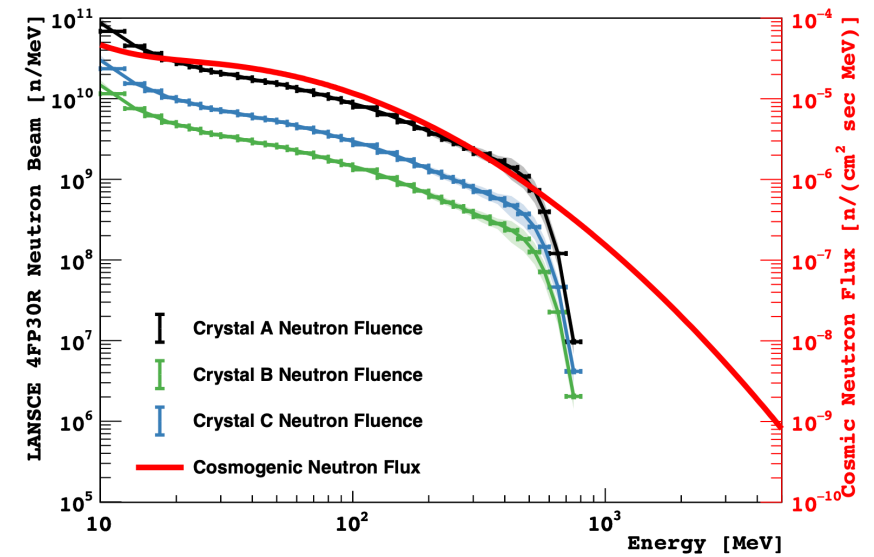
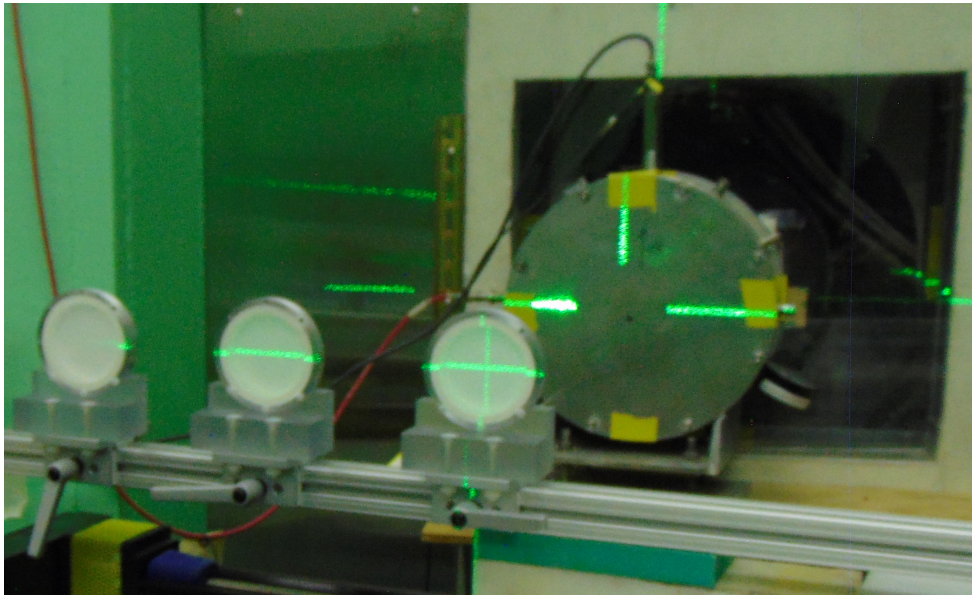


Phys. Rev. D 107, 022006 (2023)  
arXiv:2209.14898





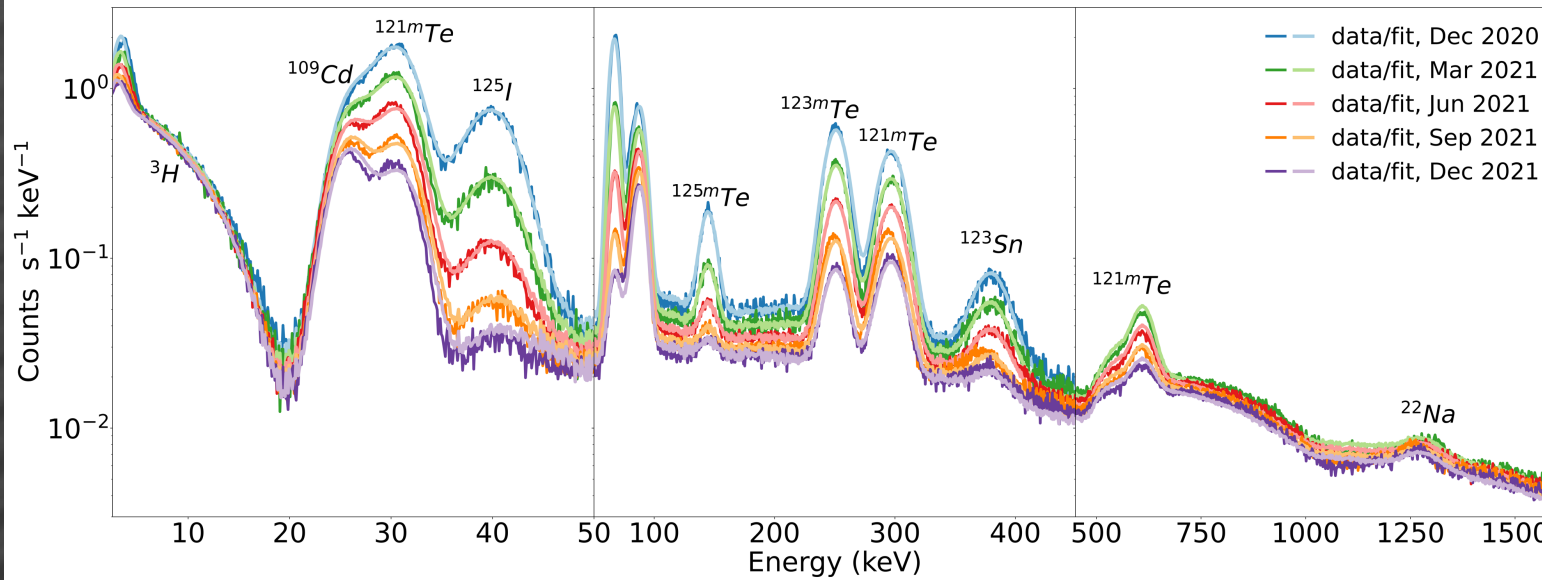
# Nal Beam Irradiation



Irradiate NaI crystals on LANSCE beam with different exposures (concerned about possible radiation damage)

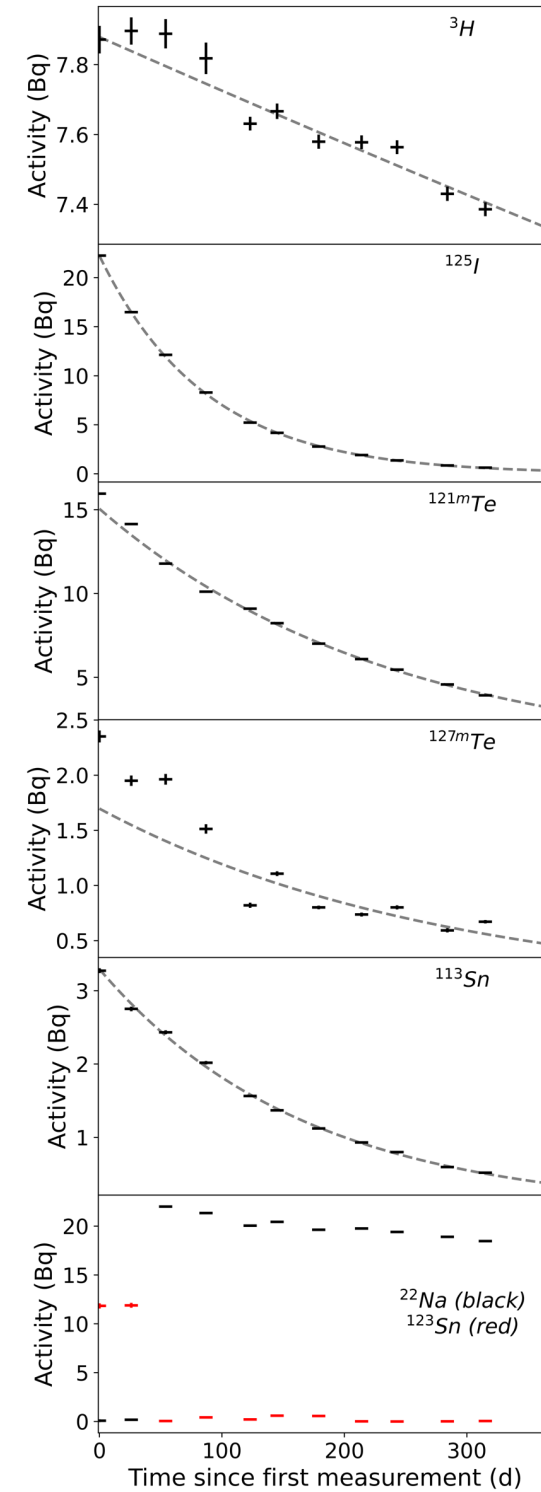


# Nal Counting

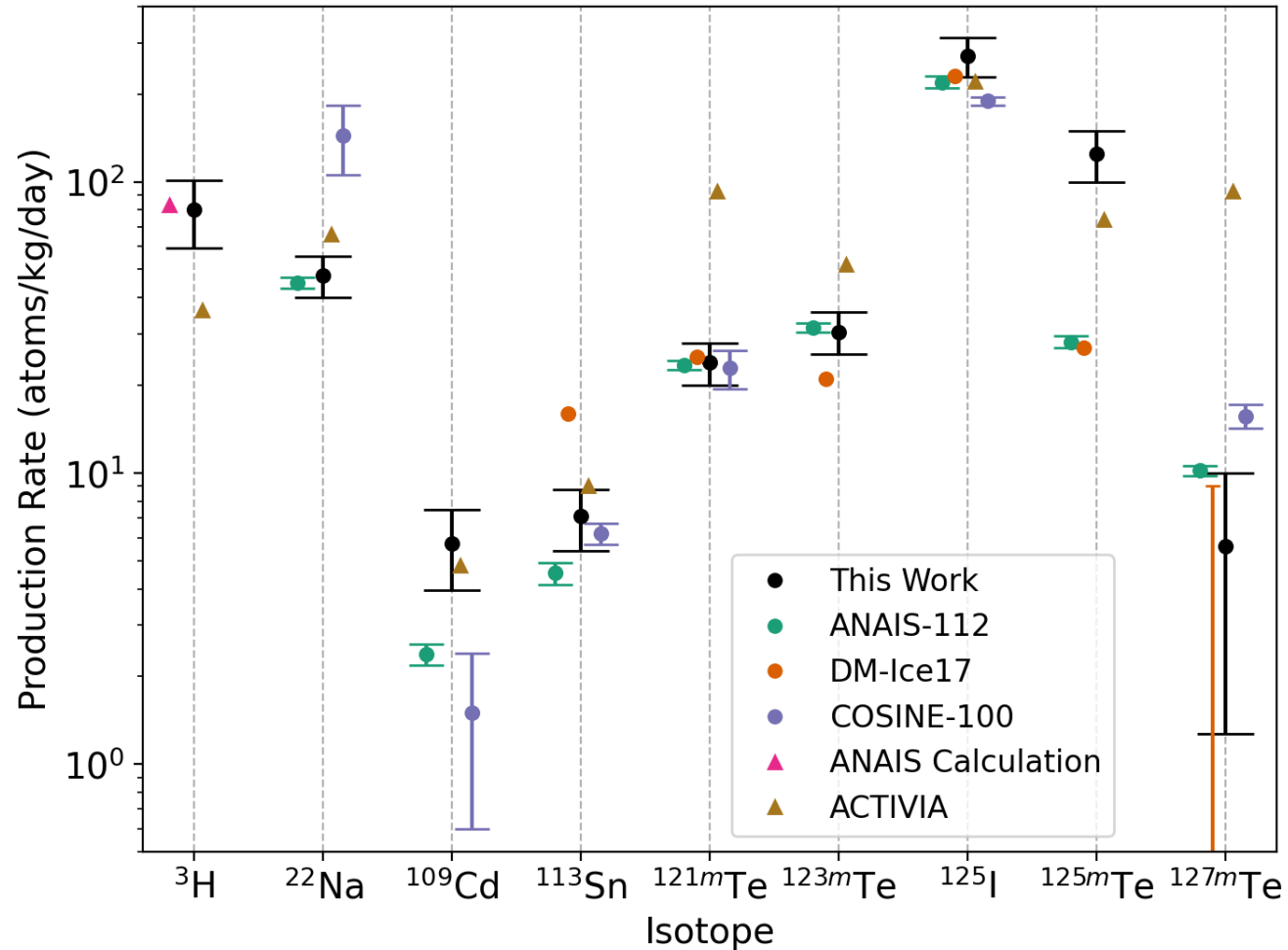


Mounted PMT to crystal and self-counted induced activity

Data was fit in both energy and time to identify rates of several different isotopes



# Production Rate Results



By extrapolating the measured activity to the sea-level cosmic neutron flux (accounting for cross-section shape uncertainties) we measured the production rates of 9 different activation products, including the first experimental measurement of tritium production in NaI

## Cosmogenic Production Rate

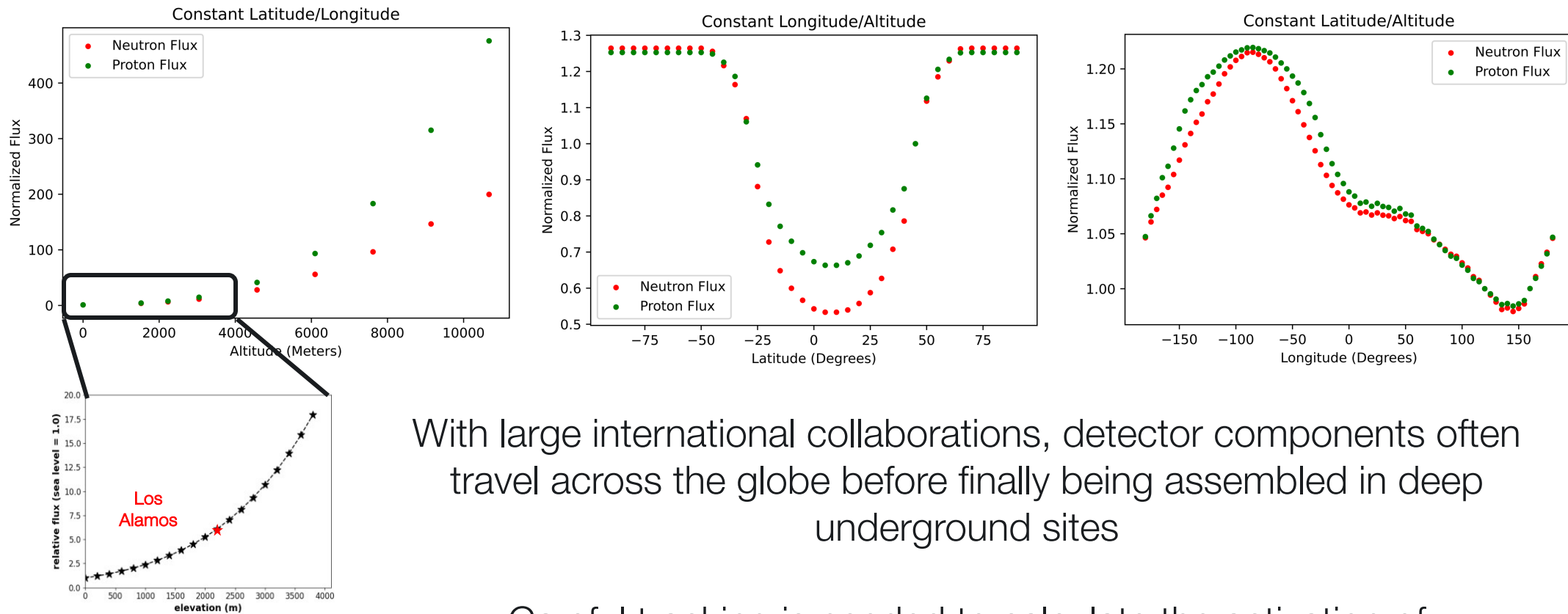
$$P(\vec{r}, t) = \int \Phi(E, \vec{r}, t) \sigma(E) n dE$$

The production rate depends on 3 things

- Cosmogenic particle flux
- Production cross-section
- Target material

# Cosmogenic Flux Variation

Flux of cosmogenic neutrons and protons varies by altitude, latitude, and longitude



With large international collaborations, detector components often travel across the globe before finally being assembled in deep underground sites

Careful tracking is needed to calculate the activation of components during transportation and storage

## Witness Samples

Witness samples: Specifically chosen materials that can be measured for induced activity to determine the cosmogenic exposure of detector parts

### Advantages:

- Direct measurement of the relevant particle flux (compared to muon flux counters)
- Can be tailored to the specific reactions of interest
- Simple, passive system, easily deployed

### Disadvantages:

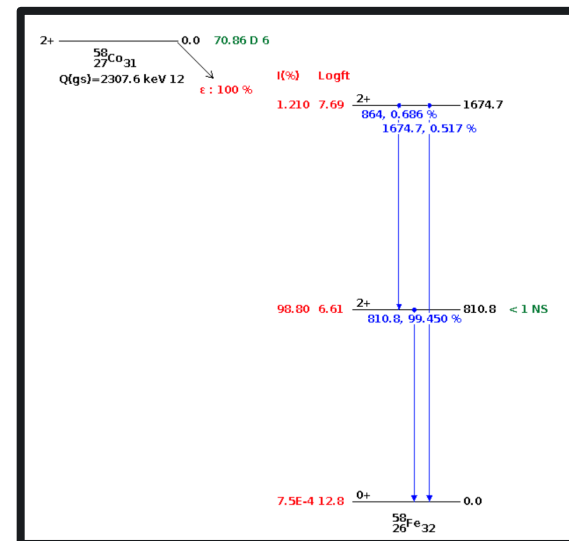
- Production rates for cosmogenic-induced reactions are low.
  - Cannot be used as a live monitor. Must be transported to a facility to be analyzed. Exposure and radioactive decay during transport must be taken into account



# Identifying Witness Materials

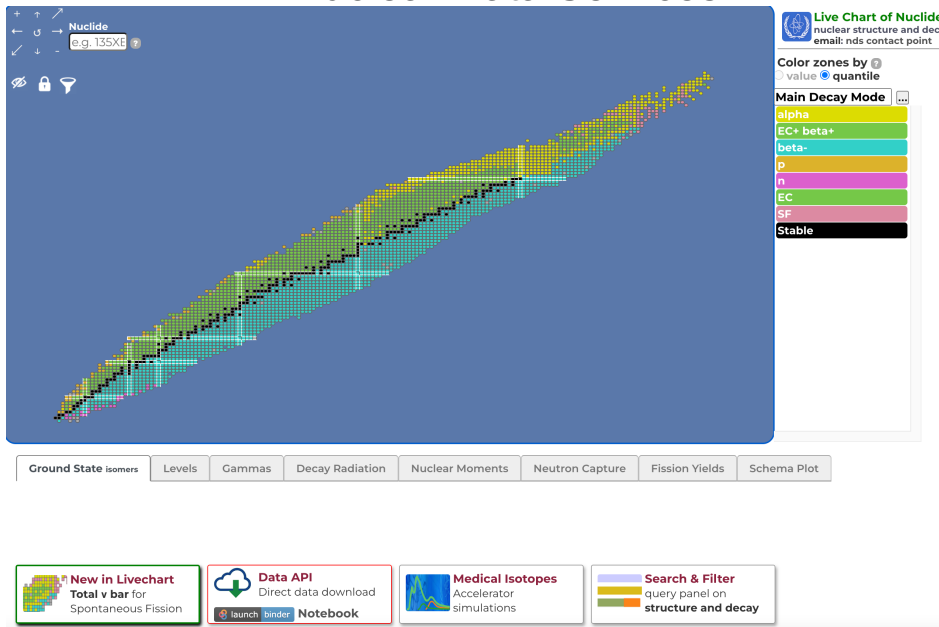
- Should contain a target isotope (with high natural abundance) with a nearby lighter radioactive isotope that
  - Has decay products that can be easily detected (e.g. gamma emitters)
  - Has a production cross-section that allows easy detection of activity (i.e. should not require a dark matter detector to be able to measure it)
  - Should have a relatively short half-life
- Ideally:
  - Reaction will be particle-type specific (e.g. (n,p), or (p,n))
  - Target will have single element composition
  - Target will be non-toxic, relatively cheap, and easy to procure.

<b>58Ni</b> STABLE 68.0769%	<b>59Ni</b> 8.10e+4 y $\epsilon + \beta + = 100\%$	<b>60Ni</b> STABLE 26.2231%
(n,p)		
<b>57Co</b> 271.8 d $\epsilon = 100\%$	<b>58Co</b> 70.883 d $\epsilon + \beta + = 100\%$	<b>59Co</b> STABLE 100%
(n,2n)		
<b>56Fe</b> STABLE 91.754%	<b>57Fe</b> STABLE 2.119%	<b>58Fe</b> STABLE 0.282%

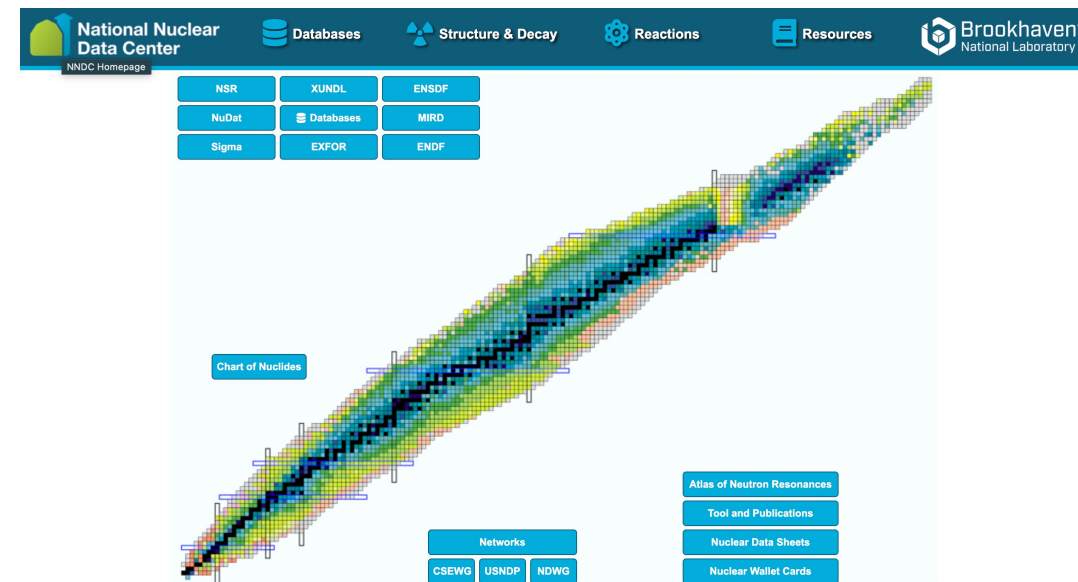


# Identifying Witness Materials

IAEA Nuclear Data Services

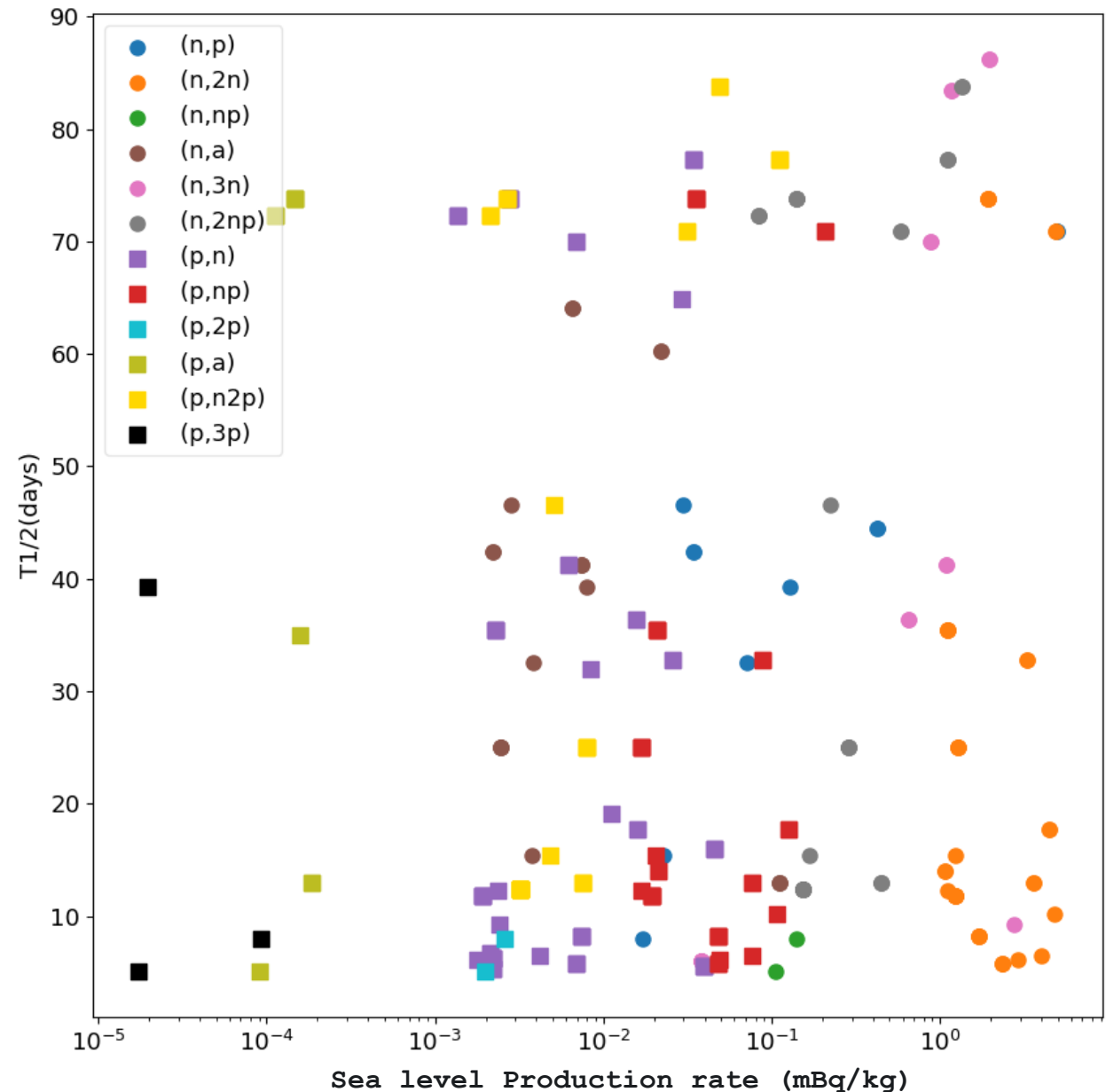
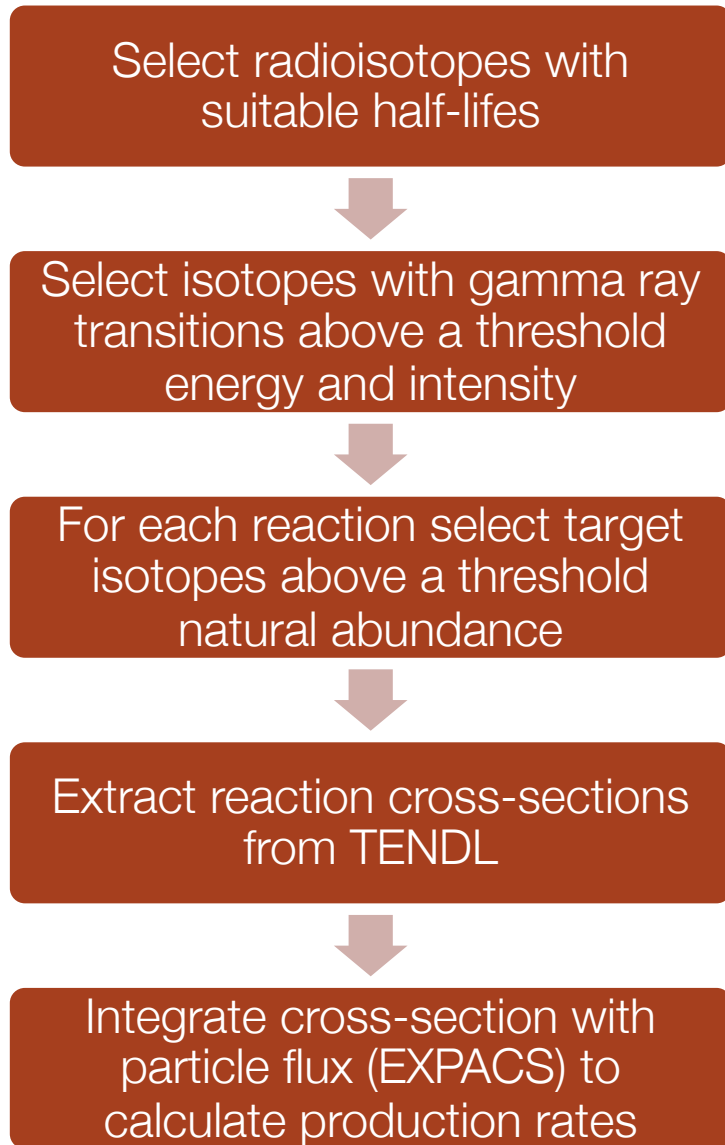


BNL National Nuclear Data Center



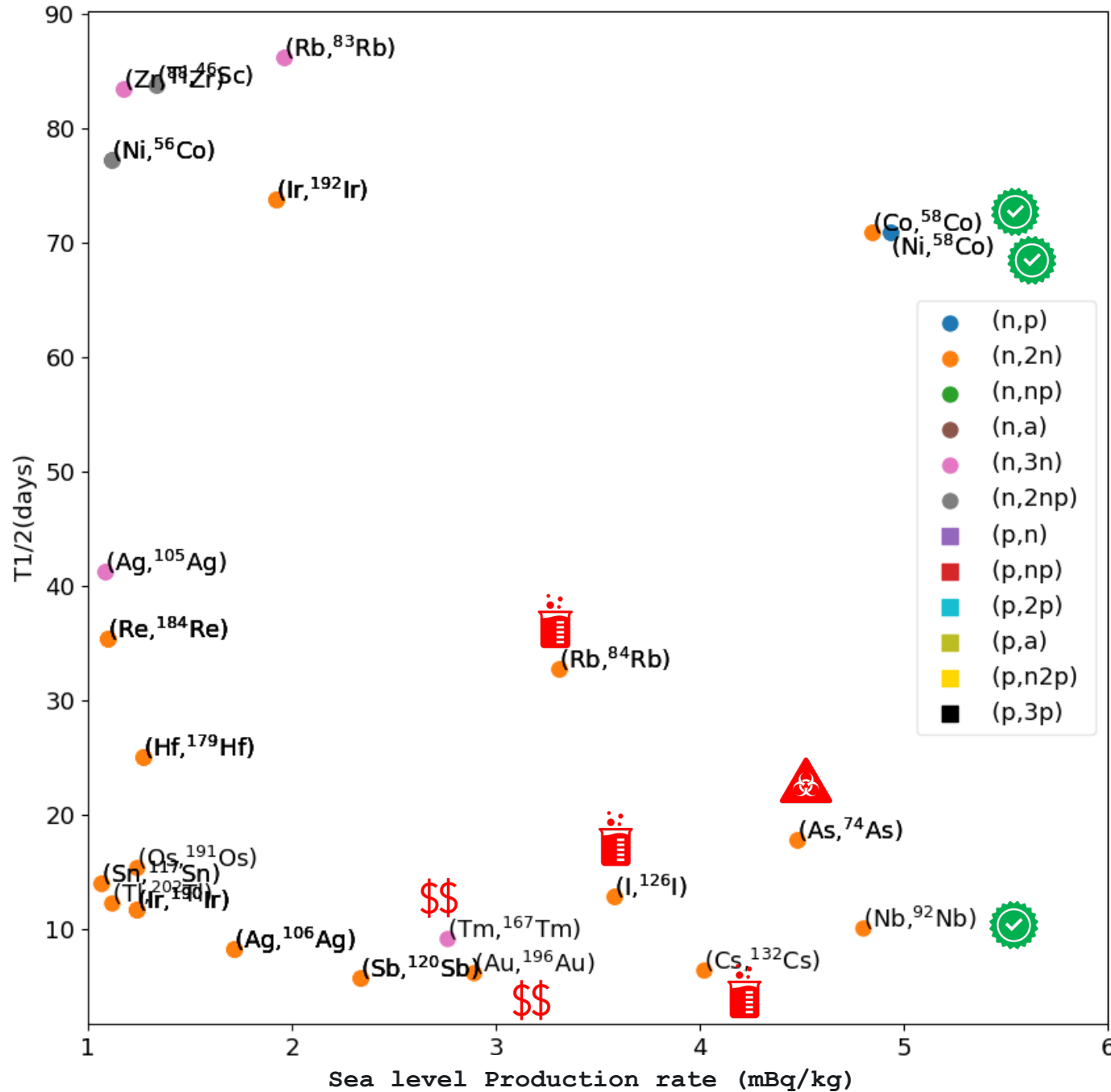
Use existing isotope databases to collect full set of possible isotope for witness materials  
 Downloaded all relevant isotope data and saved in dataframe for subsequent work  
 ~ 1600 isotopes  
 ~ 190,000 decay schemes

# Identifying Witness Materials



(n,2n) reactions typically have highest production rates  
 Proton-induced reactions having lower rates due to the lower flux at sea-level

# Candidate Witness Materials



Besides production rates and half-lives, one also needs to take into account toxicity, chemical reactivity, price, etc.

# Deploying Witness Materials

Target	Reaction	Product	Product Half-life [days]	Time to 95% Saturation [days]	PNNL Sat. Activity [mBq/kg]	LANL Sat. Activity [mBq/kg]
Ni	(n,p), (n,2np), (n,3np)	<sup>58</sup> Co	70.9	307	5.5	31.5
Co	(n,2n), (p,np)	<sup>58</sup> Co	70.9	307	5.6	32.4
Nb	(n,2n)	<sup>92m</sup> Nb	10.2	44.1	1.93	11.2
Ti	(p,n)	<sup>48</sup> V	16.0	69.2	0.057	0.39

Aim to target locations and paths that explore largest flux variations

- Sea-level, High-altitude ground
- Low, High Latitude
- Cross-Pacific/Atlantic Flight

Witness materials will be packaged with GPS tracker to evaluate activation during transportation

## Goals

- Establish feasibility of using witness materials to measure high energy (> 10 MeV) particle flux
- Develop software for community to track activation based on location and materials



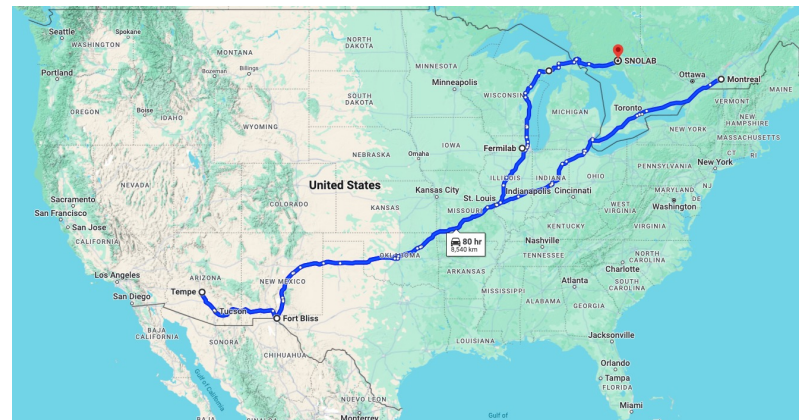
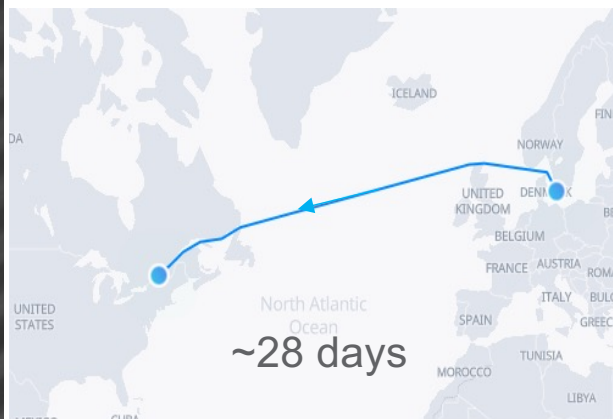


# Cosmogenic Tritium

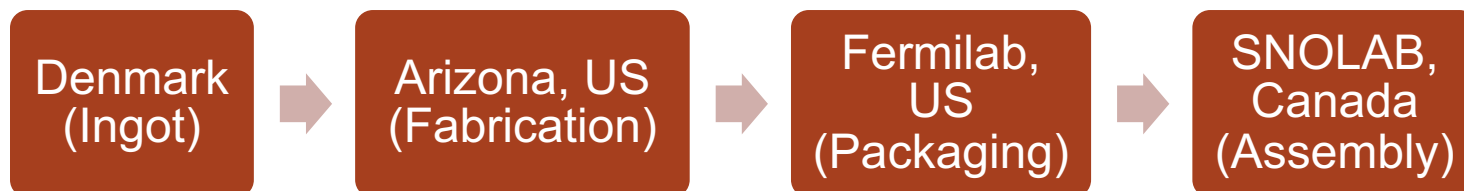
Sea-level production rate of tritium was measured to be  
 $124 \pm 25$  atoms/kg/day

Corresponds to a background rate of  
 $\sim 1.8$  mDRU per day of sea-level exposure in the 0-2 keV energy range

Target Background Goal for Oscura CCD Dark Matter Experiment:  
**25 mDRU**



Extremely short window of time to transport silicon from Europe to North America and fabricate the CCDs, even accounting for shielded transportation and underground packaging.

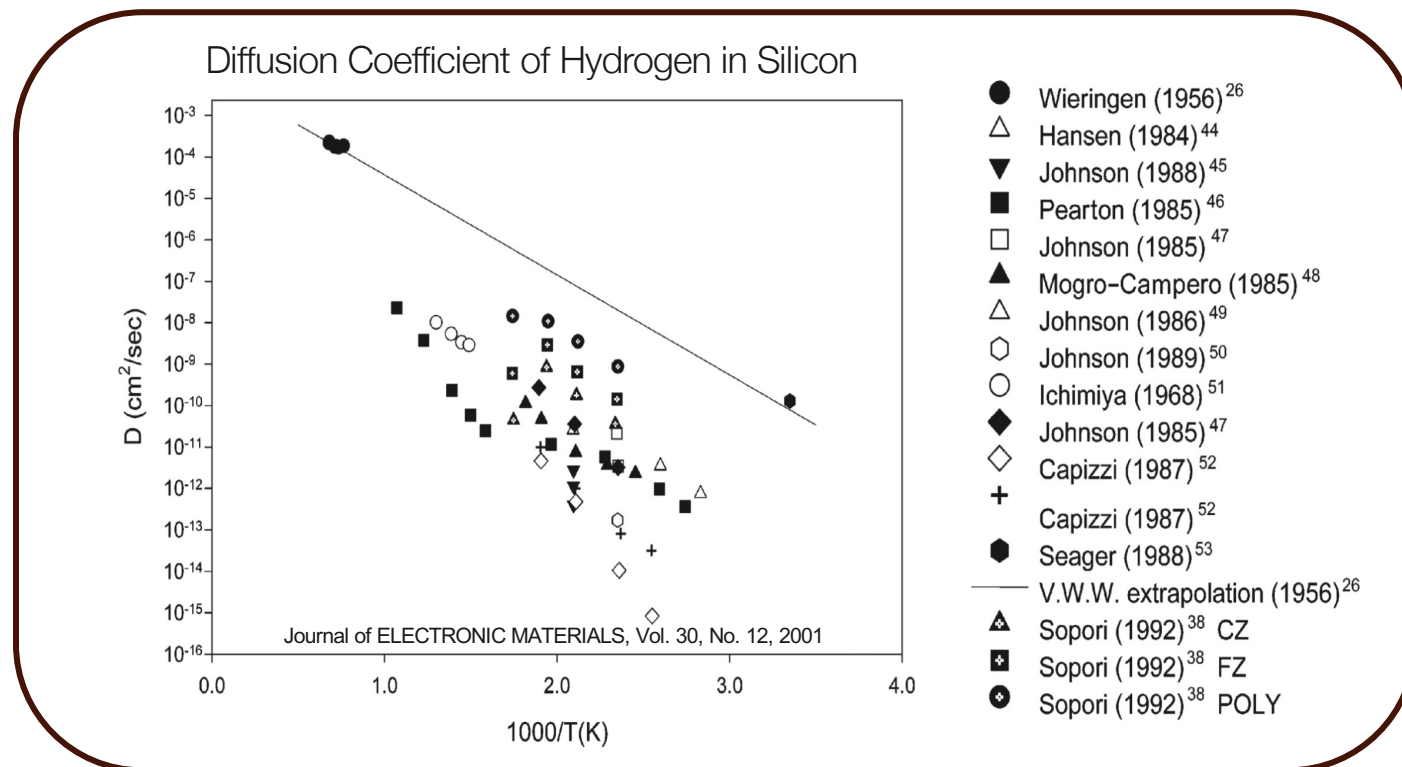


Is there another solution?

# Tritium Bakeout

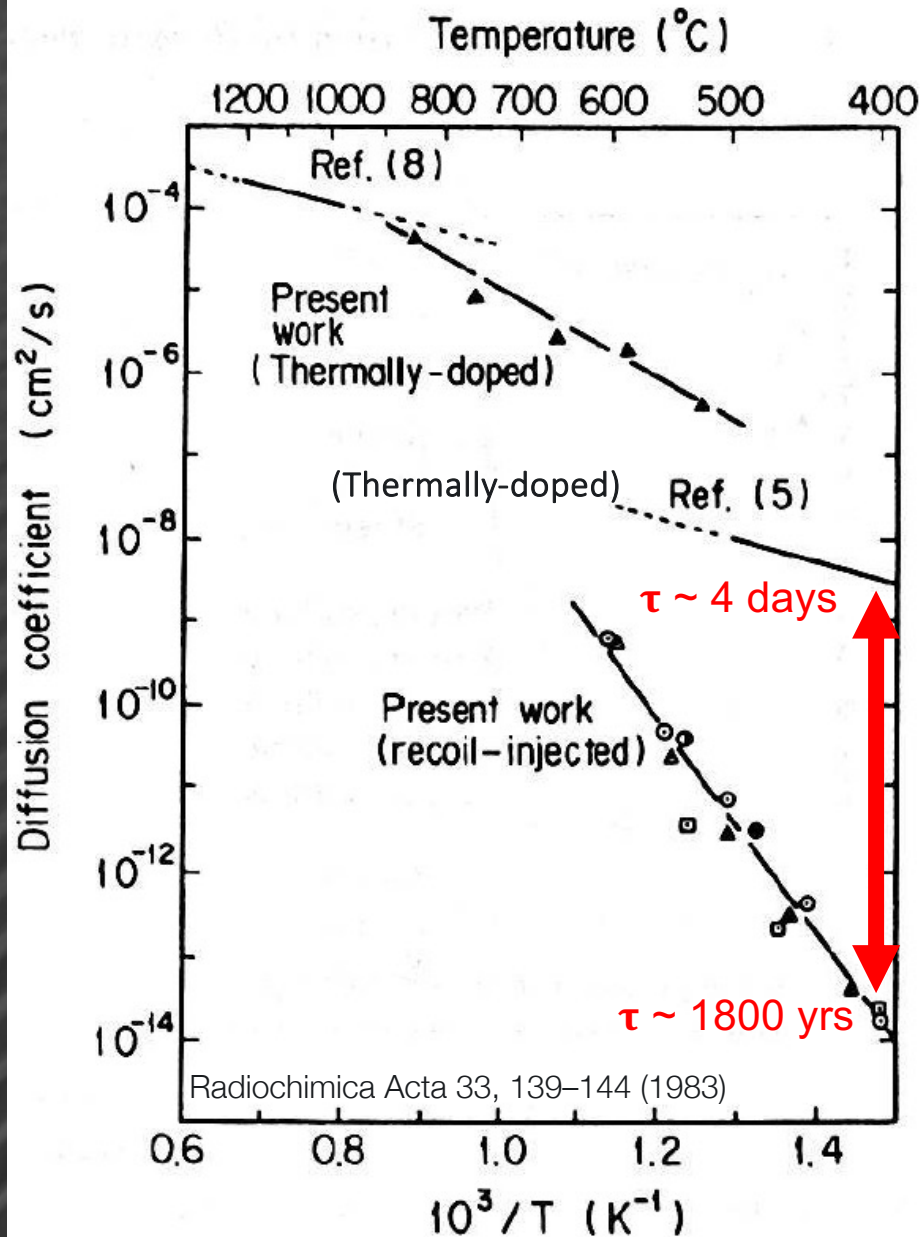
Hydrogen is typically mobile as an interstitial species in the silicon lattice

Diffusion of hydrogen in silicon is commonly used in semiconductor industry to reduce recombination and improve device characteristics.



Can we remove cosmogenic tritium from silicon by baking at elevated temperatures?

# Tritium Bakeout



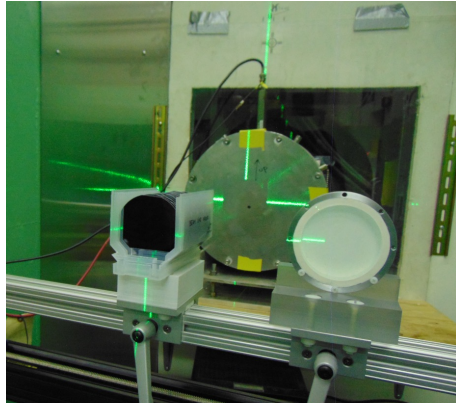
For tritium in silicon, the diffusion coefficient seems to strongly depend on whether the tritium was thermally diffused or implanted through nuclear recoils

This large difference is attributed to possible damage of the silicon lattice during recoil implantation, leading to trapping of tritium on defects

Since cosmogenic tritium is produced through spallation of a silicon nucleus, it is likely to also produce damage and defects in the silicon lattice

# Tritium Removal Experiment

Silicon wafers on  
LANSCE neutron beam



RADDEC Pyrolyser-Trio Furnace  
System



Quantulus Ultra Low  
Background Liquid Scintillation  
Counter



Irradiate  
on  
neutron  
beam



Produce  
tritium  
through  
spallation



Bake silicon  
at elevated  
temperatures



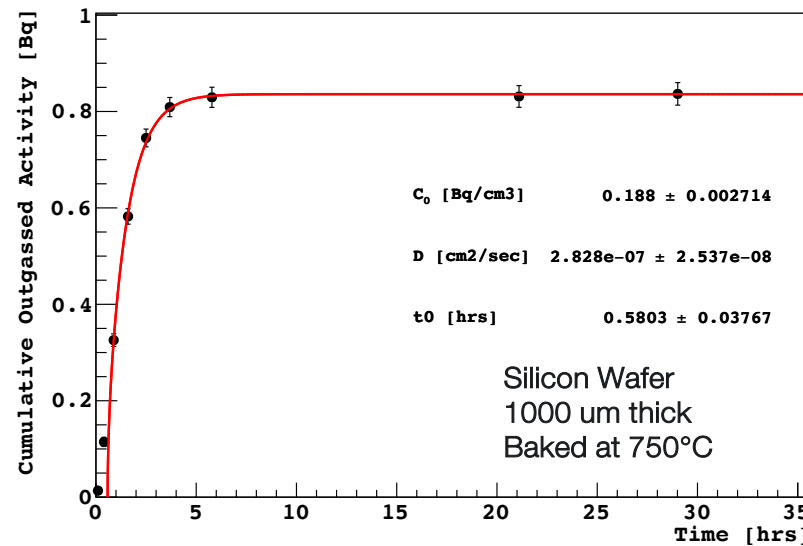
Collect  
removed  
tritium as  
a function  
of time



Count  
tritium  
activity in  
liquid  
scintillator



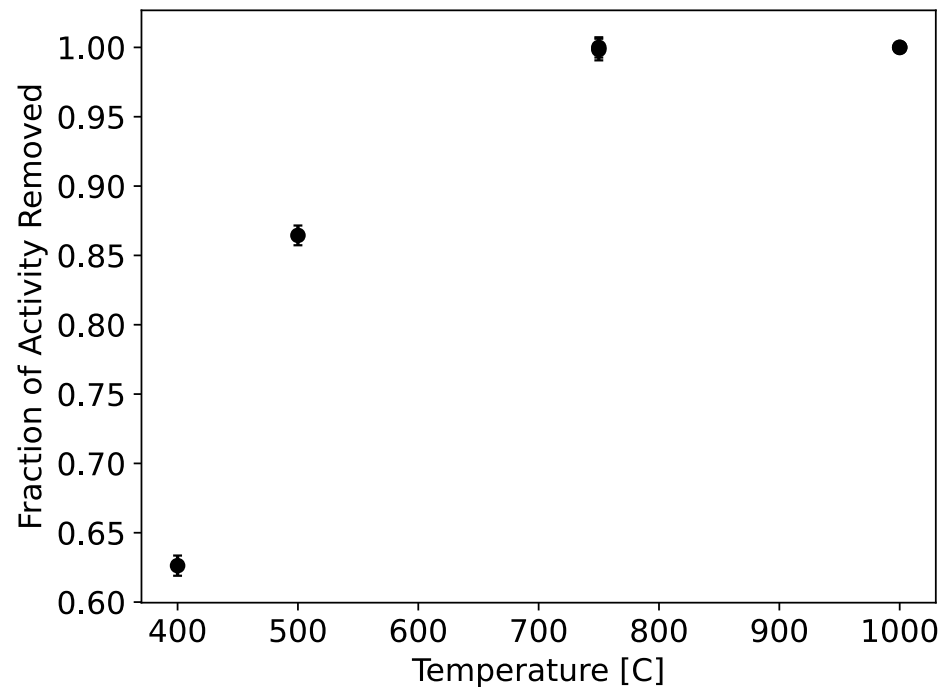
Evaluate  
diffusion  
and  
removal  
fraction



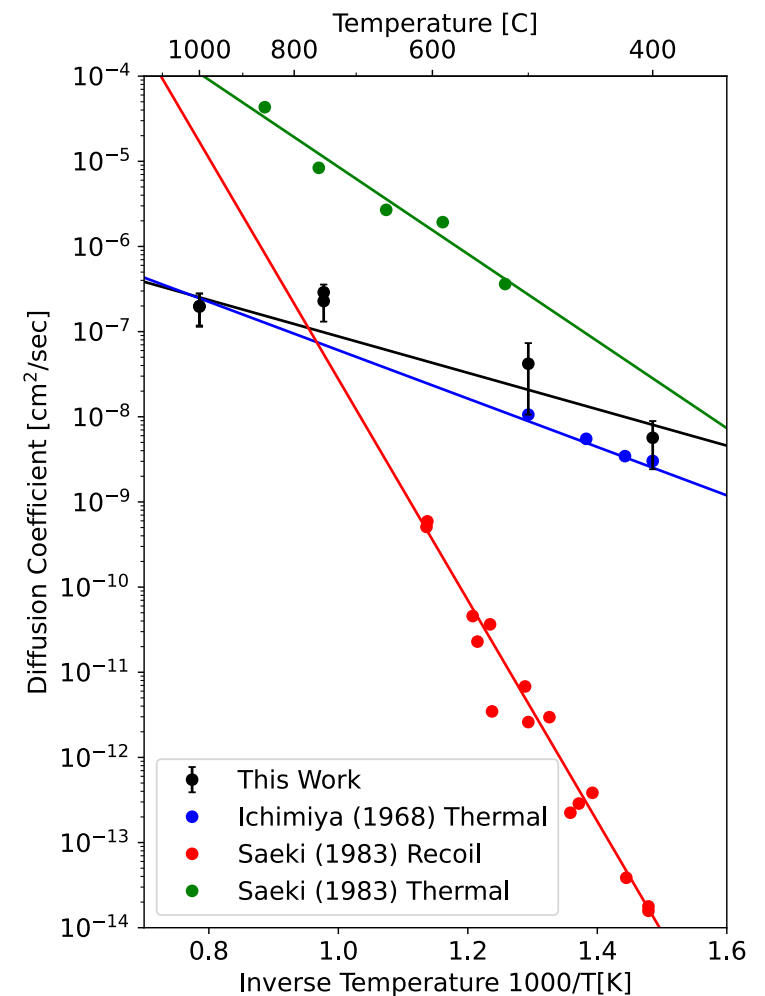
# Tritium Removal Experiment

**PRELIMINARY**

Spallation-induced tritium can be completely removed from high-purity silicon with baking above 750°C for a few hours and diffusion coefficients are consistent with thermally diffused measurements



Results imply that cosmogenic tritium created prior to device fabrication can be removed by a high temperature bake, greatly reducing the constraints on exposure and shielding





## Summary

- Understanding the production rates of cosmogenic backgrounds is critical for next generation rare event searches
- Production rates can be measured using the neutron beam at LANSCE and other locations. Measurements have been in argon, silicon, and sodium iodide
- High energy particle flux can be measured using activation of carefully selected witness materials
- In certain cases, removal techniques may be possible.  
We have demonstrated the ability to remove spallation-induced tritium from silicon using elevated temperatures.



# Thank You

## Si Activation



**Los Alamos National Laboratory**  
S.R. Elliott

**Pacific Northwest National Laboratory**  
R. Bunker, J. Burnett, R. Saldanha\*, R. Tsang

**University of Chicago**  
A. Matalon, P. Privitera, K. Ramanathan, R. Thomas

**University of Washington**  
A. Chavarria, P. Mitra, A. Piers

## Nal Activation



**Australian National University**  
L. Bignell, G. Lane, Y. Zhong

**Los Alamos National Laboratory**  
S.R. Elliott

**Pacific Northwest National Laboratory**  
R. Saldanha\*, R. Tsang, M. Zalavadia

**University of Illinois at Urbana Champaign**  
L. Yang

**Yale University**  
R. Maruyama, W. Thompson

## Tritium Removal



**GAU, University of Southampton**  
D. Reading\*, P. Warwick

**Pacific Northwest National Laboratory**  
R. Saldanha\*, B. Loer

**University of Chicago**  
P. Privitera

**University of Washington**  
P. Mitra, A. Chavarria

## Witness Samples



**Pacific Northwest National Laboratory**  
A. Hellinger, R. Saldanha

**Los Alamos National Laboratory**  
S.R. Elliott, R. Massarczyk



U.S. DEPARTMENT OF  
**ENERGY**

Office of  
Science