

# Optimised of the neutron yield calculation from $(\alpha, n)$ reactions with modified SOURCES4 code

M. Parvu<sup>1</sup>, P. Krawczun<sup>2</sup> and V. A. Kudryavtsev<sup>2</sup>

<sup>1</sup> University of Bucharest, Romania
<sup>2</sup> University of Sheffield, United Kingdom

## Outline

- Introduction: neutron production in  $(\alpha, n)$  reactions in SOURCES4
- 'Optimised' cross-sections and comparison with data.
  - Cross-checks with alpha beam data.
- Neutron yields and spectra from 'optimised' SOURCES4 versus data for radioactive decay chains.
- Conclusions

See my talk at LRT2024 (and contribution to the proceedings) for previous update.

Preprint available: Parvu et al., arXiv:2408.10910 (2024)

# **SOURCES4**

- SOURCES4A/4C: W.B. Wilson, et al., SOURCES4A: a code for calculating (*α*,*n*), spontaneous fission, and delayed neutron sources and spectra, Technical Report LA-13639-MS, Los Alamos, 1999;
- Working historically with SOURCES4A; no noticeable difference for our goals.
- The probability for an alpha particle to produce a neutron by interacting with a nuclide i ( $N_i$  is the number density of atoms of nuclide i):

$$P(E_{\alpha}) = \int_{0}^{E_{\alpha}} \frac{N_{i}\sigma_{i}(E)}{\left(-\frac{dE}{dx}\right)} dE$$

- Stopping power cross-sections from tables compiled by Ziegler.
- Approximation of thick target.

## **Recent modifications to SOURCES4A**

- The **maximum number of discrete nuclear levels** for the product nuclides was increased from 100 to **500**.
- The **maximum number of target elements** was increased from 20 to **110**.
- The cross-sections for (α, n) reactions in SOURCES4 have been taken from reliable experimental data (including some recent ones) where possible and complemented by the calculations with TALYS1.96, EMPIRE2.19/3.2.3, and JENDL-5 where the data were scarce or unavailable.
- Various sets of cross-sections are available in the library → we just recommend the most reliable in our opinion.
- The code was modified so the user does not need to change the order of the cross-section or branching ratios anymore, but only to indicate which one to use (the order that the cross-section appears in the library) → data selection option in tape1.

Previous modifications: Carson et al. Astropart. Phys., 21 (2004) 667; Lemrani et al. NIMA 560 (2006) 454; Tomasello et al. NIMA, 595 (2008) 431; Tomasello et al. Astropart. Phys., 34 (2010) 70, Kudryavtsev et al. NIMA 972 (2020) 164095; Kudryavtsev et al. AIP Conf. Proc. 2908 (2023) 1, 100003.

## **Advantages and disadvantages**

- Advantages
  - Flexible libraries of cross-sections and branching ratios
  - Fast calculation
  - Total neutron spectra; spectra from interactions on individual isotopes and from the variety of radioisotopes in a single calculation; spectra from the ground state and different excited states.
- Disadvantages:
  - Written long time ago (but cross-sections can be added/replaced)
  - Written in Fortran (but no need to intervene if the code works)
  - No gammas generated from de-excitation of final state nuclei (same for other codes)
  - Cannot read ENDF format (but if you know ENDF format, converting the crosssection data into the SOURCES4 format is not a big deal).
  - Cannot deal with 'surface' contaminations/problems.

# **Cross-sections for <sup>13</sup>C (low-A target)**



- Codes (based on statistical models) do not predict resonance structure of the crosssections for light elements.
- Only <sup>13</sup>C contributes to the neutron yield on carbon (fraction: 1.07%).
- Harissopulos 2005 + TALYS 1.96 calculations 'default' in SOURCES4

 $\rightarrow$  leads to higher neutron yields in materials containing carbon-

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## New measurements for <sup>13</sup>C



- New measurements from Brandenburg et al, PRC, 108 (2023) L061601.
- Not yet in SOURCES4 library.

# Oxygen



Kunz up to 1 MeV + Bair + JENDL-2021 above 5.3 MeV

JENDL-2021 (overlapping with data)

- The choice of cross-section above 5.3 MeV is not obvious.
- JENDL follows experimental data below 5.3 MeV and can be the best option to use above this energy.

## **Uranium oxide**



 JENDL-2021 above 5.3 MeV provides a better agreement with data from alpha beams than TALYS1.96.

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#### 19**F**



- Sometimes data do not allow us to choose the optimum cross-section, as for <sup>19</sup>F.
- The cross-sections measured by Peters (2016) complemented with those from TALYS1.96 above 7 MeV.
- The choice of the code above 7 MeV is not critical but the choice of the data below 7-8 MeV is important.

#### <sup>27</sup>AI



- The cross-sections measured by Flynn and Howard were used complemented with those calculated with EMPIRE 3.2.3 above 9.2 MeV.
- Not a big difference compared to models.

#### $^{10}B(\alpha,n)^{13}N$ (19.9%)



Typical example of multiple data sets not agreeing with each other both for cross-sections and neutron yields from alpha beams.

Prior + JENDL 2021 – default in SOURCES4

#### $^{11}B(\alpha,n)^{14}N$ (80.1%)



Prior + JENDL 2021 – default in SOURCES4. The choice is driven by the agreement observed with alpha beam data.

#### <sup>nat</sup>B( $\alpha$ ,n)N



B-11 has 80.1% abundance  $\rightarrow$  B-10 is not so critical.

\* Vlaskin's points are **not** experimental data, but an evaluation of neutron yield based on experimental data.

#### **Neutron spectra: 5 MeV alphas**



19F

<sup>27</sup>Al

## **Neutrons from radioactive decay chains**



- Fernandes et al. EPJ Web of Conferences, 153 (2017) 07021.
- Not the real measurements but evaluation of neutron yields from radioactive decay chains from alpha beam measurements
  - Systematic uncertainties due to the procedure used may be quite high.
- Higher yield for carbon from SOURCES4 not a surprise (but see next slide).
- Fe: Lower for U and ok for Th -> uncertainty in 'measurements'?
- The agreement is within 10% for most materials tested.

## **Neutrons from radioactive decay chains**



- G. V. Gorshkov, O. S. Tsvetkov, Soviet Atomic Energy, 14 (1964) 573–577.
- Direct measurements of neutron yields from radioactivity.
- Carbon looks fine but with quite a big uncertainty.
- Strangely, NaF shows higher neutron yield in SOURCES4 whereas CaF<sub>2</sub> better agrees with data? Still some uncertainty in data?
- The agreement is within 10% for most materials tested.

## Conclusions

- Statistical models (TALYS1.9 and EMPIRE2.19/3.2.3) are not recommended for use for light isotopes at low energies where the cross-sections show a resonant behaviour (from the authors of the codes).
- An optimised approach:
  - Use data where possible (usually at low alpha energies and no controversy) and a model at higher energies that agrees with data at low energies
  - If no data exist for an isotope, use a model (TALYS or EMPIRE) based on comparison of the neutron yields with alpha beam data (if available)
  - A model for branching ratios.
  - Comparison with alpha beam data show good agreement (<sup>13</sup>C is still a question).
- Neutron yields from decay chains:
  - Neutron yields with the optimised approach show a good agreement with data
    - within 10% for most materials (mainly light elements)
  - Calculated neutron spectra agree reasonably well with the measured ones
    - differences still exist but the measurements are not easy.

## **Backup: table with neutron yields**

| Element | $^{nat}\mathbf{U}$     | $^{232}$ Th            | Compound                   | $^{nat}U$              | $^{232}$ Th            |
|---------|------------------------|------------------------|----------------------------|------------------------|------------------------|
| Li      | $7.12 \times 10^{-10}$ | $2.95\times10^{-10}$   | $Al_2O_3$                  | $8.53 \times 10^{-11}$ | $4.17 	imes 10^{-11}$  |
| Be      | $8.38 \times 10^{-9}$  | $2.79 	imes 10^{-9}$   | BeO                        | $3.08 \times 10^{-9}$  | $1.03 \times 10^{-9}$  |
| В       | $1.99 \times 10^{-9}$  | $6.14 \times 10^{-10}$ | $C_2F_4$                   | $9.76 \times 10^{-10}$ | $3.90 \times 10^{-10}$ |
| С       | $1.76 \times 10^{-11}$ | $7.04 \times 10^{-12}$ | $CaCO_3$                   | $7.28 \times 10^{-12}$ | $2.89 \times 10^{-12}$ |
| N       | $5.80 	imes 10^{-11}$  | $3.23 \times 10^{-11}$ | $CaF_2$                    | $6.63 \times 10^{-10}$ | $2.75\times10^{-10}$   |
| Na      | $4.13 \times 10^{-10}$ | $1.93 	imes 10^{-10}$  | $CH_2$                     | $1.71 \times 10^{-11}$ | $7.04 	imes 10^{-12}$  |
| Mg      | $2.03 \times 10^{-10}$ | $7.67 \times 10^{-11}$ | $H_2O$                     | $3.98 \times 10^{-12}$ | $1.39 \times 10^{-12}$ |
| Al      | $1.67 \times 10^{-10}$ | $8.25 \times 10^{-11}$ | $H_3BO_3$                  | $3.38 \times 10^{-10}$ | $9.64 \times 10^{-11}$ |
| Si      | $2.18 \times 10^{-11}$ | $1.01 \times 10^{-11}$ | MgO                        | $1.20 \times 10^{-10}$ | $4.56 	imes 10^{-11}$  |
| Р       | $2.85 	imes 10^{-11}$  | $1.94 \times 10^{-11}$ | $Na_2CO_3$                 | $2.78 \times 10^{-10}$ | $1.29 \times 10^{-10}$ |
| Cl      | $8.08 \times 10^{-11}$ | $4.36 \times 10^{-11}$ | NaCl                       | $1.52 \times 10^{-9}$  | $6.07 \times 10^{-10}$ |
| Ar      | $1.52 \times 10^{-10}$ | $9.00 \times 10^{-11}$ | NaF                        | $8.16 \times 10^{-10}$ | $3.50 \times 10^{-10}$ |
| Ca      | $1.80 \times 10^{-12}$ | $1.22 \times 10^{-12}$ | $PbF_2$                    | $4.39 \times 10^{-10}$ | $1.74 \times 10^{-10}$ |
| Ti      | $3.63 	imes 10^{-11}$  | $3.18 	imes 10^{-11}$  | $SiO_2$                    | $1.41 \times 10^{-11}$ | $5.98 	imes 10^{-12}$  |
| Cr      | $1.40 \times 10^{-11}$ | $1.45 \times 10^{-11}$ | Stainless steel            |                        |                        |
| Mn      | $9.29 \times 10^{-12}$ | $1.04 \times 10^{-11}$ | Fe(66%), Cr(17%), Ni(12%), | $7.38 	imes 10^{-12}$  | $8.91\times10^{-12}$   |
| Fe      | $4.74 \times 10^{-12}$ | $6.68 \times 10^{-12}$ | Mn(2%), Mo(2%), Si(1%)     |                        |                        |
| Ni      | $1.02 \times 10^{-13}$ | $2.63\times10^{-13}$   | UC                         | $2.50\times10^{-12}$   | $1.02\times10^{-12}$   |
| Cu      | $3.67	imes10^{-13}$    | $1.07 	imes 10^{-12}$  | $UO_2$                     | $2.41 \times 10^{-12}$ | $8.37	imes10^{-13}$    |

Whole U and Th decay chains. Units: n/g/s/ppb.

Additionally, spontaneous fission of  ${}^{238}$ U gives  $1.35 \times 10^{-11}$  n/g/s/ppb.

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